



## Abstract

It is generally believed that subduction of lithospheric slabs is a major contribution to thermal heterogeneity in Earth's entire mantle and provides a main driving force for mantle flow. Mantle structure can, on the one hand, be inferred from plate tectonic models of subduction history and geodynamic models of mantle flow. On the other hand, seismic tomography models provide important information on mantle heterogeneity. Yet, the two kinds of models are only similar on the largest (1000s of km) scales and are quite different in their detailed structure. Here, we provide a quantitative assessment how good a fit can be currently achieved with a simple viscous flow geodynamic model. The discrepancy between geodynamic and tomography models can indicate where further model refinement could possibly yield an improved fit. Our geodynamical model is based on 300 Myr of subduction history inferred from a global plate reconstruction. Density anomalies are inserted into the upper mantle beneath subduction zones, and flow and advection of these anomalies is calculated with a spherical harmonic code for a radial viscosity structure constrained by mineral physics and surface observations. Model viscosities in the upper mantle beneath the lithosphere are  $\sim 10^{20}$  Pas, and viscosity increases to  $\sim 10^{23}$  Pas in the lower mantle above D". Comparison with tomography models is assessed in terms of correlation, both overall and as a function of depth and spherical harmonic degree. We find that, compared to previous geodynamic and tomography models, correlation is improved significantly, presumably because of improvements in both plate reconstructions and mantle flow computation. However, high correlation is still limited to lowest spherical harmonic degrees. An important ingredient to achieve high correlation – in particular at spherical harmonic degree two – is a basal chemical layer. Subduction shapes this layer into two rather stable hot but chemically dense "piles", corresponding to the Pacific and African Large Low Shear Velocity Provinces. Visual comparison along cross sections indicates that sinking speeds in the geodynamic model are somewhat too fast, and should be  $2 \pm 0.8 \text{ cm yr}^{-1}$  to achieve a better fit.

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# 1 Introduction

At convergent plate margins, slabs of subducted lithosphere start their journey toward the Earth's interior, and seismic tomography is arguably the best tool to track their sinking. Based on such seismic models, the opinion of most scientists is currently that most slabs eventually sink to the base of the mantle (Grand et al., 1997; van der Hilst et al., 1997; van der Voo et al., 1999), where they accumulate. However, slab sinking trajectories are complicated through their interaction with phase transitions, particularly the spinel-perovskite transition at 660 km depth, where some slabs may lay flat for a while before they sink further (e.g. Fukao et al., 2001). While long-debated, the interactions of slabs with the transition zone, and the degree of mass transport, are still unclear.

By comparing slab locations predicted from geodynamic models based on subduction history, both plate reconstructions (e.g. Bunge and Grand, 2000; Hafkenscheid et al., 2006; Zhang et al., 2010; Shepard et al., 2012) and geodynamic model parameters, such as slab sinking rates and mantle viscosity (e.g. Ricard et al., 1993; Lithgow-Bertelloni and Richards, 1998; Liu et al., 2008; van der Meer et al., 2010; Čížková et al., 2012), can be constrained. Towards that goal, we present here a simple geodynamic model of mantle density based on subduction history, and compare it to seismic tomography. We both visually compare along cross sections and compute formal correlations (cf. Ray and Anderson, 1994; Becker and Boschi, 2002). Our work is essentially an update of Steinberger (2000), which we believe is appropriate now, as both models of seismic tomography and subduction history have changed since then. We also compare our results with those of a simple slab sinker approach (Ricard et al., 1993; Lithgow-Bertelloni and Richards, 1998), as well as an updated slab sinker approach based on our own subduction history model and the sinking rate  $1.2 \text{ cm yr}^{-1}$  inferred from van der Meer et al. (2010), in order to assess whether the geodynamic model in fact leads to an improvement.

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Our geodynamic model also includes a thermochemical layer in the lowermost mantle, which is shaped by subduction into piles (e.g. McNamara and Zhong, 2005; Garnero and McNamara, 2008; Zhang et al., 2010). A similar global comparison has recently also been performed by Shephard et al. (2012). However, they use a different mantle convection code and mantle viscosity structure, their mantle model does not include thermochemical piles, they use a somewhat different subduction input model, and they also compare their results to a different tomography model.

## 2 Methods

### 2.1 Geodynamic models

The geodynamic models are very similar to Steinberger and Torsvik (2012) – mostly, unless noted otherwise, the thermo-chemical model, as shown in their Fig. 5 – where the emphasis was on the creation of plumes, whereas here we focus on subduction. Since Steinberger and Torsvik (2012) describe their model in more detail, we give here only a brief description: it is based on 300 Myr of subduction history in the true polar wander (TPW) corrected, global hybrid reference frame (Steinberger and Torsvik, 2008; Torsvik et al., 2008) corrected in longitude according to van der Meer et al. (2010) (Fig. 1), i.e. density anomalies are added beneath the locations of subduction at each time step. The differences between Steinberger and Torsvik (2012) and the models st12den-1 and st12den-2 shown in this paper are the depth dependence of thermal density anomalies (for both models), where and when density anomalies are added (for st12den-2), and the consideration of phase boundaries (for st12den-2): we use here a stronger depth dependence of thermal expansivity  $\alpha$  than Steinberger and Torsvik (2012): we assume a radial profile of  $\alpha$  corresponding to the lower mantle profile of Steinberger and Calderwood (2006) but extrapolated to the surface. Consequently, density anomalies corresponding to subducted slabs become considerably smaller as they sink through the lower mantle. In model st12den-2 we add density anomalies

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Flow is modeled with a spherical harmonic approach (Hager and O’Connell, 1979, 1981). We use radial viscosity profile M2b of Steinberger and Calderwood (2006) – that is, the shape of the profile in each layer (upper mantle, transition zone, lower mantle) is inferred from mineral physics, and absolute viscosity values in each layer are adjusted to optimize fit to geoid and global heat flux (for a flow model based on seismic tomography) and satisfy the “Haskell” constraint from postglacial rebound (Mitrovica, 1996). We use a tangential stress free surface as upper and lower boundary condition, include diffusion of heat through the CMB, but apply an “isolating” upper surface, in order to not include twice the same effect, which is already considered by the explicit addition of subducted slabs into the model. Our model also includes a chemical layer at the base of the mantle, which is initially 70 km thick at the time of model initiation at 300 Ma.

As our model output, we compute the present-day temperature or thermal density field, which is compared with a seismic tomography model as discussed in the next section. As a model enhancement, we have introduced slab tracers. They are also added vertically below subduction zones at depths 650 and 700 km and advected with the flow using a 4th order Runge-Kutta scheme (Press et al., 1986). They carry as information their weight, location and time of insertion. In this way, we can more directly compute sinking times and speeds at specific locations, as well as averages and variability of sinking times and speeds and lateral motion, and compare these modeling results with observation-based estimates (e.g. van der Meer et al., 2010) to complement our own comparison with tomography.

### 2.2 Tomography model

A great number of tomography models have been published over recent years. Here we mainly compare to the SMEAN composite S-wave tomography model of Becker and Boschi (2002) (Fig. 2) which was obtained by RMS-weighted averaging of S20RTS (Ritsema and van Heijst, 2000), NGRAND (a 2001 update of Grand et al., 1997), and SB4L18 (Masters et al., 2000). SMEAN has been found to outperform other models

in geodynamic tests (Steinberger and Calderwood, 2006) and yields good variance reductions when put to the test with actual seismic waveforms (Qin et al., 2009).

We also constructed an update, which we call SMEAN2, where we replaced S20RTS by S40RTS (Ritsema et al., 2011) and NGRAND by TX2008 (the purely seismic model of Simmons et al., 2009). Differences between SMEAN2 and SMEAN are rather minor: correlations are close to unity up to spherical harmonic degrees  $l \sim 14$ , and average correlation up to degree  $l = 20$  is 0.95, and we also find that different mean models – such as a simple average of these three models, or the average of Buitert et al. (2012), or similar averages where we only include those models that have global coverage, or only S-wave models with global coverage, look all very similar. Correlation and visual agreement among different S-wave models – and even more so among the various mean models constructed – is generally much higher than between tomographic and geodynamic models (which will be discussed below; cf. Becker and Boschi, 2002).

### 2.3 Ways of comparison between geodynamic and tomographic models

We assume here that S-wave velocity anomalies are linearly related to the density anomalies we infer from our slab models, implying that all chemical heterogeneity is captured by the chemical piles in the lower mantle. This assumption ignores the potential complexities which may arise from mineral physics and thermodynamic considerations of the shear wave to density anomaly scaling in a heterogeneous mantle (e.g. Ricard et al., 2005; Stixrude and Lithgow-Bertelloni, 2010).

We compare geodynamic and seismic models both visually along cross sections and in map view and formally in terms of correlation coefficients. Formal correlation is computed and displayed in the same way as in Becker and Boschi (2002): that is, we plot or give global correlation,

- as a function of depth and spherical harmonic degree;
- as a function of depth for all spherical harmonics up to  $l = 8$  or  $l = 20$ ;

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the mantle, older slab particles will obviously move horizontal above the CMB, and distances get larger.

## 3.2 Visual comparison between geodynamic and tomography models

### 3.2.1 Map views of slabs in the entire lower mantle

Figure 4 shows the distribution of slabs predicted by our forward model and displayed in analogous way to the tomography model in Fig. 2. On the positive side, we note that there is an overall agreement between regions where slabs occur in both models. In particular, the agreement is quite good in the lowermost mantle (violet colors). Of course, this does not come as a surprise, as slabs in our model do not move very much laterally until they reach the lowermost mantle (see Sect. 3.1), and the agreement of subduction zone locations through geologic history (Fig. 1) and tomography of the lowermost mantle (Fig. 2) has been noted early on (e.g. Richards and Engenbretson, 1992). On the downside, the maps look quite different on a smaller scale: the tomography model looks more “blobby” whereas the geodynamic model generally shows linear anomalies often connected to surface subduction zones and getting more diffuse and less strong further down in the mantle, before spreading out above the core-mantle boundary. It is also not straightforward to “match” individual slabs in both models. Part of the problem is that we have to choose a certain cutoff in both figures, and the figures change with cutoff, which is why we also compare the models along cross sections in Sect. 3.2.2. But it is also clear that trying to match slabs inferred from subduction history and tomography requires a very dedicated effort (e.g. van der Meer et al., 2010).

The dependence of fit on wavelength will be more formally discussed in Sect. 3.3; expectedly, the agreement is looking much better if you step back and don’t look at details.

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### 3.2.2 Cross sections

While map views as presented in Sect. 3.2.1 enable an overall comparison, the problem of having to choose a certain cutoff, and the fact that slabs higher up in the mantle hide slabs beneath makes cross sections more suitable for a detailed comparison trying to match slabs in both models. Such a comparison of cross sections through the mantle beneath subduction zones of the past 300 Myr is given in Fig. 5. While the cross sections beneath the most of the Americas, Australia and Antarctica mostly cross one subduction zone for a given time, those beneath Indonesia, Eurasia and Alaska typically cross two – one in the circum-Pacific Ring and one corresponding to Tethys, Mediterranean and Indian Ocean subduction at the southern margin of Eurasia. Cross sections are shown for model st12den-2. Cross section for s12den-1 have more diffuse slabs – as the slab input is distributed over a larger depth interval – but otherwise look similar.

Tomographic and geodynamic cross sections show overall agreement on the large scale, with – in most cases – the central parts of the cross sections being dominated by cold or seismically fast material that can be attributed to subduction, while on one (240–285°), or both (90–225°) of the sides often hot or seismically slow material appears, due to upwellings that may occur along the margins of the Large Low Shear Velocity Provinces (LLSVPs; Garnero and McNamara, 2008). A different pattern occurs under Asia, with subduction and corresponding slab anomalies (Circum-Pacific and Tethys) on either side, and a slow anomaly (seen in the tomography cross section especially at 345°, in the model cross section at 345° and 330°) in between. A slow anomaly in the lowermost mantle beneath Russia and Kazakhstan appears in most recent tomography models and seems to be a robust feature. The tomography cross section at 345° and qualitative agreement with modeling results (see also Steinberger and Torsvik, 2012) indicate that it may well be overlain by a mantle plume.

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A more detailed comparison of slow regions, except for those corresponding to the LLSVPs of the lowermost mantle (see Sect. 3.1.3) is not attempted here, for the following reasons:

- a. Our code does not consider lateral viscosity variations, and therefore our resulting upwellings are probably unrealistically wide.
- b. A statistical comparison of seismically slow regions and predicted plume conduits from geodynamic modeling has already been shown to display good agreement for some deep-rooted plumes being connected to hotspots (Boschi et al., 2007, 2008).
- c. Although we generally find a pattern of plumes along LLSVP margins (Steinberger and Torsvik, 2012) their locations do not exactly correspond to observed hotspots.
- d. Individual, thermal plumes may be hard to detect seismically (e.g. Hwang et al., 2011).

If we attempt to match individual features in the seismically fast regions, we find they are generally less deep in the tomography model. We begin the comparison under South America ( $135\text{--}90^\circ$ ) where it is perhaps most straightforward, and then move counterclockwise around the Pacific before discussing subduction at the southern margin of Eurasia: at  $135^\circ$ , we find a gap (corresponding to smaller amounts of subduction) at a radius  $\sim 0.7\text{--}0.75$  (normalized to Earth's radius) in the geodynamic model, whereas a similar gap occurs in the tomography model at about  $0.73\text{--}0.85$ . At  $120^\circ$ , this gap occurs at larger depth in both cases:  $\sim 0.6\text{--}0.7$  in the geodynamic vs.  $\sim 0.63\text{--}0.8$  in the tomographic model. Further north, we observe more or less continuous fast or cold material from the top to bottom of the mantle in both cases. However, maxima (corresponding to largest amounts of subduction) occur again at larger depths  $\sim 0.8$  in the geodynamic model vs.  $\sim 0.85$  in tomography, at  $105\text{--}75^\circ$ . We also find that modeled maxima typically occur  $\sim 10^\circ$  further west than seismically observed, which may,

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at least in the southern part, be caused by flat slab subduction (Isacks, 1988) not accounted for in our model, where we assume vertical sinking.

Beneath North America (15–60°), disagreement becomes more prominent: still we find maxima generally deeper (~0.65–0.7) in the geodynamic model than in the tomography model (~0.72–0.8) and further to the west. Again, flat slab subduction (Bird, 1988) could, at least partly, be responsible for this lateral offset. Accordingly, Bunge and Grand (2000) invoked low-angle subduction from a comparison of geodynamic modeling results and tomography. It has been shown to be difficult to achieve agreement between geodynamic and tomographic models in this region without ad-hoc assumptions such as a stress guide (Liu et al., 2008). However, beneath North America also the shape of the anomaly disagrees. Whereas tomography shows a slab dipping from west to east (especially in the sections at 60° and 45° in Fig. 5), at an approximately constant dip angle, hence indicating an approximately constant slab sinking speed of ~1–1.5 cm yr<sup>-1</sup> (Grand et al., 1997), the modeled slab has a more complicated shape, reaching a maximum depth for slabs of about 80 Ma, whereas older slabs are less deep in the mantle.

Qualitatively, we can understand this shape due to high rates of subduction beneath North America in our model in the late Mesozoic (darker colors in Fig. 1). These not only cause relatively fast slab sinking speeds, but also an upward return flow to the side of it, further enhanced by active upwelling (indicated by red colors beneath the slab in the cross sections) thus hindering the older slabs from sinking further or even pushing them up again. However, the observation that this shape is not seen may indicate that rates of Farallon subduction beneath North America before ~80 Ma were higher than in our model. This could be the case, as the absolute Pacific plate motion is not well known before the age of the Hawaii and Louisville hotspot chains, and hence Circum-Pacific rates of subduction are not well constrained before ~80 Ma. Larger amounts of Farallon subduction could also account for the discrepancy between seismically fast material at the base of the mantle at 60° and 45° beneath North America, and the absence of corresponding material in the geodynamic model.

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this could be because sinking speeds are too fast in the model, but it has also been suggested that an ancient slab is being drawn up beneath the discordance (Gurnis and Müller, 2003).

In the regions further south, across and near Antarctica (cross sections at 210° and 180°) agreement is good again to the extent that in both models the most prominent fast or cold anomalies only occur in the lowermost mantle, because subduction has terminated in that region at about 80 Ma. At the 180° (south polar) cross section it can also be seen that, once subduction has stopped, the last slabs sink at a considerably slower rate than average shown in Fig. 2: from the computed present-day depth ~ 1300 km and the insertion depth 650–700 km at 66 Ma an average sinking speed of ~ 1 cm yr<sup>-1</sup> is inferred. Given that the tomographic anomaly reaches up to about the same depth, one can infer a similar sinking speed also from tomography.

Slabs subducted at the southern margin of Eurasia can be seen at the right side of the cross sections from 30° to 0° and at the left side from 345° to 270°. We can again mostly match fast anomalies in the seismic with cold anomalies in the geodynamic cross sections. Again, anomalies in the geodynamic cross sections occur often (such as in the 315° cross section) at greater depth than in the seismic ones. There is no strong lateral offset, except at 270°. In that equatorial cross section beneath Indonesia, the fit is very poor: the geodynamic model predicts cold anomalies throughout the lower mantle, whereas the seismic model shows a strong anomaly laterally displaced in the upper mantle and upper part of the lower mantle. This misfit is probably in part because the plate kinematic history in that region is very complicated and not adequately matched by our simplified model.

Finally, cross sections beneath Asia, especially at 330°, show in their central parts the remains of the Mongol-Okhotsk subduction zone (van der Voo et al., 1999). Like in the other case where subduction has subsequently stopped – beneath Antarctica – the geodynamic model predicts here comparatively slow sinking of the final slabs subducted in that region, such that the geodynamic model predicts cold anomalies at a similar depth to the observed fast seismic anomalies.

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occur to the fact that upwellings in the model with piles are generated at locations that are less different from regions of low seismic wavespeed.

In contrast, if we include neither thermo-chemical piles nor diffusion of heat across the bottom thermal boundary layer in the model (case st12den-7 in Fig. 7), we obtain an even higher correlation except in the lowermost mantle: including diffusion of heat across the CMB improves correlation in the lowermost mantle, because model slabs push hot material toward locations corresponding to LLSVPs. However, higher up in the mantle, correlation gets worse, as upwellings in the model form at locations that generally do not match well with actual upwellings.

We find that correlations have improved compared to the earlier slab model of Steinberger (2000), which gives  $r_8 = 0.3$  and  $r_{20} = 0.21$  with SMEAN (Becker and Boschi, 2002). In this case, like for our new slabs-only model st12den-7 (top right panel of Fig. 7), correlation remains at a similar level throughout the mantle; and is slightly higher in the upper part of the lower mantle than at its base. Correlations of SMEAN with the simple slab sinker model (vertical sinking at a prescribed speed) of Lithgow-Bertelloni and Richards (1998) are also similar to correlations with Steinberger (2000) – ( $r_8 = 0.33$  and  $r_{20} = 0.18$ ). For comparison we have also devised a slab sinker model based on our own subduction model, both in the TPW corrected global hybrid reference frame (Steinberger and Torsvik, 2008; Torsvik et al., 2008) and in a reference frame additionally corrected in longitude (van der Meer et al., 2010) shown in Fig. 1, and vertical sinking of  $1.2 \text{ cm yr}^{-1}$  (van der Meer et al., 2010). We find that despite the update in plate reconstruction model, correlations of the slab sinker model with tomography remain low, on a similar level to the model of Lithgow-Bertelloni and Richards (1998). The fact that the “slabs-only” geodynamic model st12den-7 gives much higher correlations than the slab sinker approach emphasizes the importance of mantle flow modeling which predicts (although rather small) lateral advection of slabs, and variable slab sinking speeds (Fig. 3).

Finally, because it appears that including upwellings that dynamically form in our model always deteriorates correlation, we consider a model where, instead, we include

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plumes with surface positions based on hotspots, and tilted plume conduits with moving source at the CMB, as in Boschi et al. (2007, 2008), based on the modeling approach of Steinberger and Antretter (2006). This approach is different in that here plumes are treated as essentially passive, not influencing large-scale flow. We find that in this case (bottom right panel of Fig. 7) correlations are further improved compared to the slab-only model. We note that here the amplitude scaling of plumes and hence the amplitude of the combined model is somewhat arbitrary, but resulting correlations depend on this scaling only slightly. Note, though, that the flow field used to advect plumes in this approach is not based on subduction, but inferred from tomography.

## 4 Discussion

Slab sinking speeds in our model (Sect. 3.1) are significantly higher, by about a factor 2, than the estimate  $1.2 \pm 0.3 \text{ cm yr}^{-1}$  of van der Meer et al. (2010) based on comparing reconstructed subduction zones with tomography. Our own comparison to tomography in Sect. 3.2 essentially confirms that model sinking speeds appear to be too high. Building upon this comparison we can give our own estimate of what would be appropriate slab sinking speeds to best explain tomography: we identify characteristic features that can be visually matched in the geodynamic model and tomography model cross sections. Based on the slab tracers, we determine the age of slabs corresponding to this feature. We then obtain our own observation-based sinking speed estimate by dividing the depth of the feature in the tomography model through this age.

We distinguish between the following three cases: (a) beginning of subduction, or substantial increase in the amount of subduction (e.g. beneath South America); (b) end of subduction (especially Mongol-Okhotsk subduction beneath Asia; Phoenix subduction beneath Antarctica); and (c) specific features in the middle of subduction (the bend in the slab beneath Japan). Results are plotted in Fig. 8. It appears that most data points plot around a straight line through the origin with a slope of about  $2 \text{ cm yr}^{-1}$ , but with considerable spread, with most data points falling between lines with slope

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1.2 cm yr<sup>-1</sup> and 2.8 cm yr<sup>-1</sup>. Exceptions are the two data points corresponding to the Mongol-Okhotsk slab, which would correspond to a much lower sinking speed: however, here our geodynamic model predicts an inverted age progression, with the oldest slabs on top, as subduction at two sides of it – at the southern and eastern margins of Asia – has pushed this slab up again. On the other hand the “225” data point (Tonga-Kermadec) corresponds to sinking speed higher than 2.8 cm yr<sup>-1</sup>, which could well be caused by the fast convergence rate and corresponding large amount of subducted slab per time and subduction zone length in this region.

The discrepancy of our observation-based sinking-speed estimate with the 1.2 cm yr<sup>-1</sup> determined by van der Meer et al. (2010) is therefore somewhat marginal. We also note that our approach is somewhat biased toward high sinking speeds – at least when interpreting the lower end of a slab, as often slabs are bent in our geodynamic model, such that not the oldest slab is at the lower end.

Our modeled sinking speeds are higher than found by Čížková et al. (2012) for the case where they use a similar viscosity structure. We think that this difference occurs mainly because Čížková et al. (2012) model relatively short episodes of subduction whereas our model typically has subduction in the same region for a long time, leading to larger amounts of subducted slabs, and hence faster sinking.

The models of Shephard et al. (2012) are more similar to ours in that respect, as they are also based on actual subduction history. They find that sinking speeds in the lower mantle do not exceed 1.5–2 cm yr<sup>-1</sup>, but this difference is probably due to their viscosity being 10<sup>23</sup> Pas throughout the lower mantle, whereas in our model, such high viscosities are only reached in the lower part of the lower mantle. We find that our correlations between models and tomography are mostly higher than those of Shephard et al. (2012), in their case with 200 Myr of subduction in the “subduction reference frame”, which is their case most similar to our model. This is in part due to our comparison to a different tomography model (for example, we find somewhat lower correlations, if we compare our model with S40RTS rather than SMEAN) and, more importantly, because

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we have included a thermo-chemical layer that is formed into “piles” at the base of the mantle.

Figure 6 suggests, at least to us, that the thermo-chemical model – with two large hot regions approximately corresponding to the two LLSVPs – fits tomography better than the purely thermal one. In addition to the two LLSVPs, the tomography model shows one smaller low shear velocity province beneath north of the Caspian Sea. This feature occurs in many recent tomography models, and therefore appears to be robust. One might consider that if this feature – which is similar in size and location to one of the small hot anomalies in the thermal model – is resolved, tomography would generally resolve a pattern such as in the thermal model in the center panel of Fig. 6. However, more appropriately, the geodynamic models should also be “looked at” through a tomography filter (Megnin et al., 1997; Schuberth et al., 2009; Bull et al., 2009).

Our model provides a significant improvement compared to earlier models – both based on a simple slab sinker approach (Lithgow-Bertelloni and Richards, 1998) and on mantle flow models (Steinberger, 2000). The remaining, and still quite substantial, misfit can help to find out how the model needs to be improved in order to come closer to the ultimate goal of a subduction-based model of mantle evolution that accurately explains present-day tomography.

Firstly, we have tried to match certain features in the tomography model with corresponding features in the geodynamic model, and find that they occur in the geodynamic model generally too deep in the mantle. This might be compensated by assuming an even higher viscosity in the lower mantle. However, an even higher viscosity globally would presumably be difficult to reconcile with geoid constraints (e.g. Steinberger and Calderwood, 2006). This points towards a possibly important effect of lateral viscosity variations, which are not included in our model: if slabs have been subducted for a long time in a certain region, they cooled the lower mantle, leading to increased lower mantle viscosities in that region and thus slower slab sinking speed, while the global average

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viscosity could remain compatible with geoid constraints (Yoshida and Nakakuki, 2009; Ghosh et al., 2010).

Another difference is that in our model, the slabs often appear bent (e.g. in the cross sections beneath North America), such that sometimes older slabs are less deep than younger ones. In contrast, tomography in that region has been interpreted such that slabs sink at approximately constant speed such that a subduction zone migrating at a constant speed could give a slab with constant dip (Grand et al., 1997). Again, this discrepancy could possibly be due to lateral viscosity variations causing slabs to be stiffer than their surroundings: a stiffer slab would be less readily bent.

Although in general our model slabs are too deep, the opposite case also occurs, namely beneath the Tonga-Kermadec subduction zone. This can be attributed to the fact that the convergence rate in our model is too low, and hence illustrates the importance of considering detailed, regional plate reconstructions. Also, a cross section where the fit is particularly poor is beneath Indonesia, which is also known to be a region of particularly complicated plate tectonic history. Clearly, it would be beyond the scope of any single paper to address this problem globally, so improvements here should be made region by region, possibly still within a global model, but with regionally refined plate reconstructions and focusing on a regional comparison.

Besides including more detailed plate reconstructions, it will also be a key issue to start with plate reconstructions further back in time. Our tentative model, where we re-ran starting from the modeled present-day structure for another 300 Myr yielded about 25 % slower slab sinking speeds. We think this occurs because of the accumulation of cold slab material in the lower mantle beneath subduction zones, leading to a reduced density contrast of newly subducted slabs and the surrounding lower mantle, and thus a reduced sinking speed, even without considering lateral viscosity variations. Including plate reconstructions further back in time could hence yield a similar effect. Furthermore, it could help to keep thermo-chemical piles stable for a longer time: observational evidence indicates that they have been in a stable position since 200 Myr (Torsvik et al., 2006) and possibly much longer (Torsvik et al., 2010) whereas they form

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typically move only a few degrees laterally. However, such lateral advection and the lateral variation of slab sinking speeds are relevant, and they lead to an improvement in model fit to tomography compared to models with slabs sinking vertically at constant speed. If a chemical layer is included in the model, it yields two thermo-chemical piles in the lowermost mantle, similar in shape and location to the Large Low Shear Velocity Provinces that are seen in tomography. This model correlates very well with the SMEAN composite tomography model up to spherical harmonic degree  $\sim 3$ –4.

Comparison along cross sections shows substantial differences between geodynamic and tomographic model, but allows to match certain “slab” features in either model with each other.

Corresponding features in the geodynamic model appear normally at greater depth than in the tomography model, indicating that modeled sinking speeds are too fast. Through such matching of features, we can obtain an observation-based slab sinking speed estimate of  $\sim 2 \text{ cm yr}^{-1}$ , varying mostly between  $1.2 \text{ cm yr}^{-1}$  and  $2.8 \text{ cm yr}^{-1}$ .

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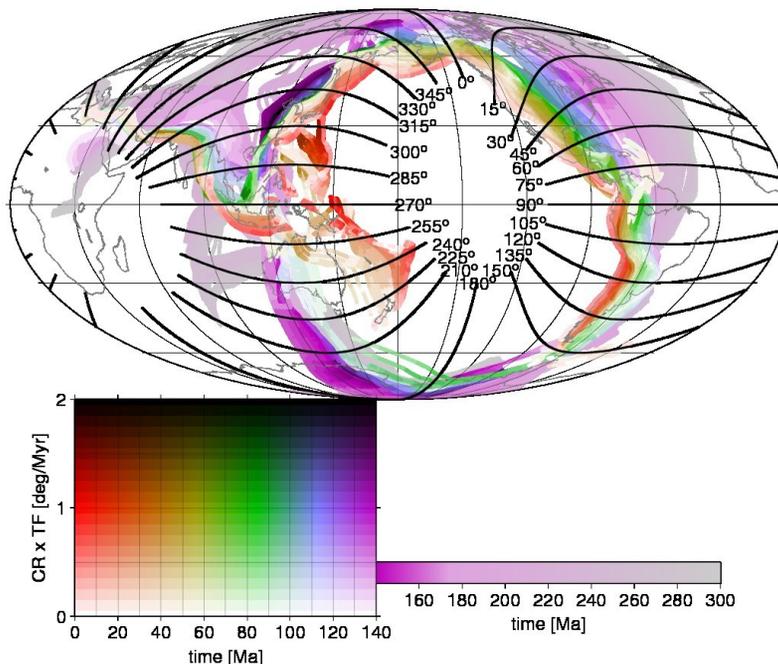
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**Fig. 1.** Subduction history and locations of cross sections shown. The subduction history model is the same as in Steinberger and Torsvik (2012), but displayed in a different and hopefully intuitive way: color represents time of subduction before present, and darkness represents the amount of subducted material per time and per subduction zone length, expressed in terms of convergence rate (CR) times a thickness factor (TF), which accounts for the increase of lithosphere thickness with age. We use  $TF = (\text{age}/80 \text{ Myr})^{1/2}$  for age  $< 80 \text{ Myr}$  and  $TF = 1$  for age  $> 80 \text{ Myr}$ . Convergence rates are largely unconstrained before 140 Ma and therefore not considered – see Steinberger and Torsvik (2010) for details. Younger slabs are plotted on top of older ones, corresponding to a “view from above” for slabs sinking vertically at constant speed (cf. Sigloch, 2011).

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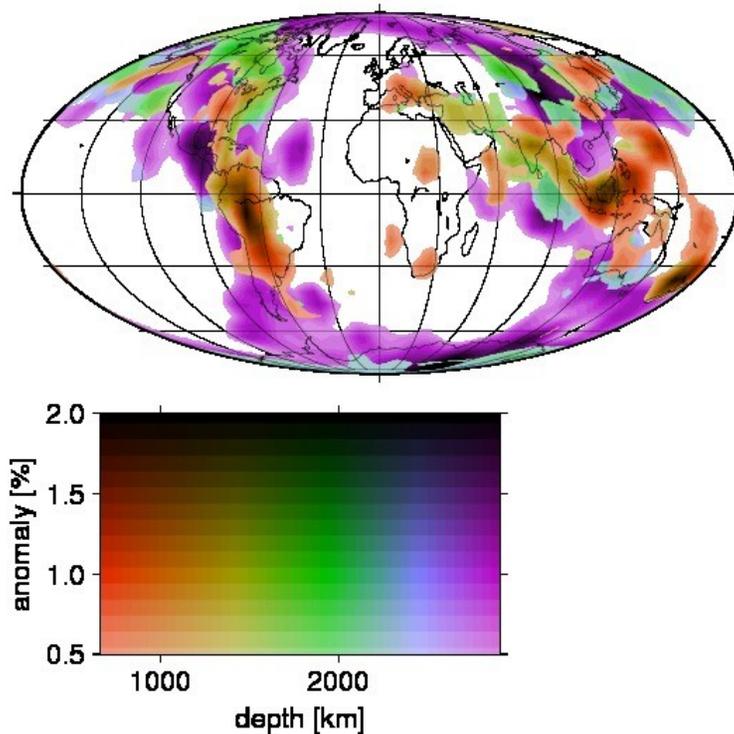
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**Fig. 2.** A representation of the SMEAN (Becker and Boschi, 2002) composite tomography model that is meant to show the depth and intensity of slab-related anomalies. For each location, we determine local maxima of the S-wavespeed anomaly vs. depth profile, as such maxima may correspond to the centers of subducted slabs. We plot the depth (represented by color) and S-wavespeed anomaly (represented by darkness) of such maxima, if they occur in the lower mantle (depth > 670 km) and exceed 0.5 %. If at a given location more than one maximum satisfying these conditions is found, only the upper one or uppermost one is plotted, corresponding to a “view from above”.

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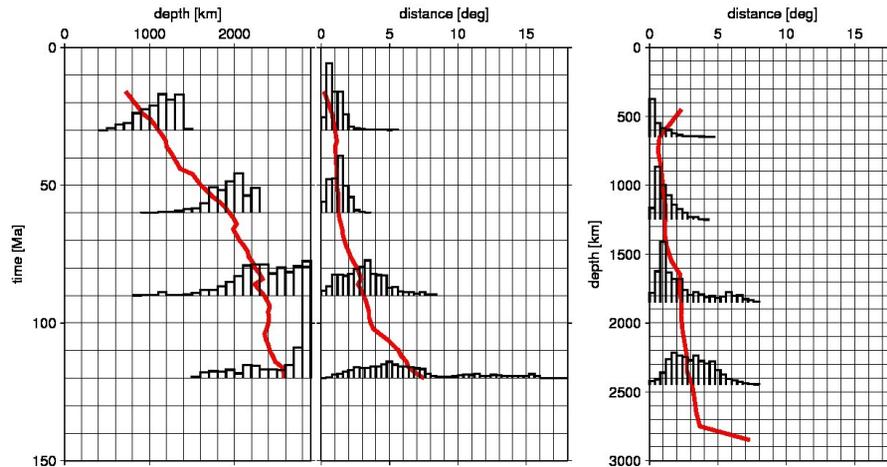
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**Fig. 3.** Left panel: depth of slab tracers vs. time of subduction. Center panel: lateral displacement of slab tracers vs. time of subduction. Right panel: lateral displacement vs. depth. All panels are for model st12den-2; red lines indicate the average, whereas histograms plotted for certain depths or times illustrate variability. Only those slabs subducted since 120 Ma are plotted, since older slabs tend to move laterally in the lowermost mantle, and may get heated up and rise again, and would hence make the picture less clear. Both average and histograms consider variable “mass” of slab tracers, which may differ because of variable convergence rate, age of subducted plate and spacing of tracers along subduction zones (which is kept nearly constant).

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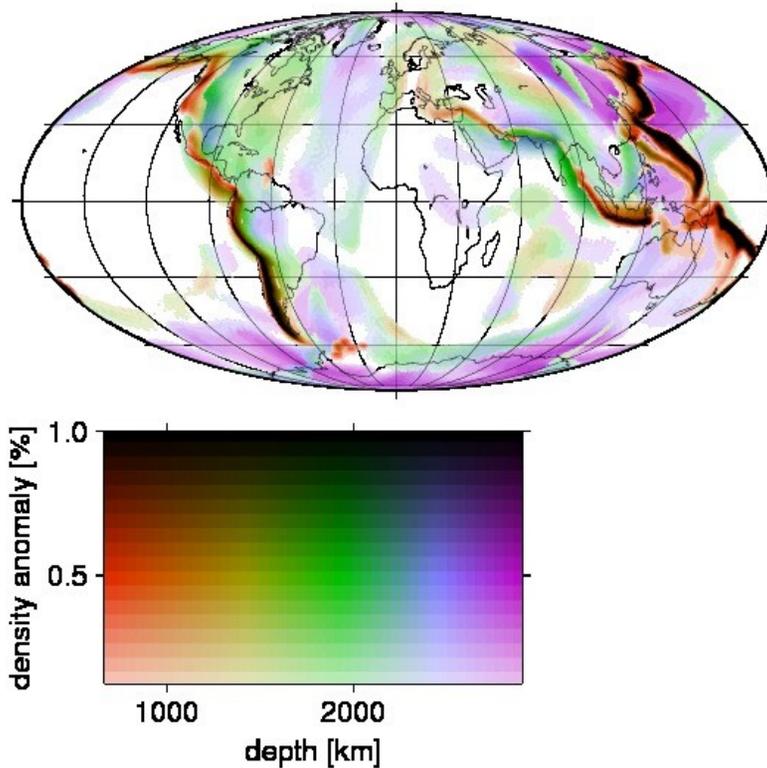
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**Fig. 4.** Depth and intensity of slabs as inferred from our geodynamic forward model st12den-2, plotted in an analogous way as the tomography model in Fig. 2: for each location, we determine local maxima of the density anomaly vs. depth profile, as such maxima correspond to the centers of subducted slabs in our model. We plot the depth (represented by color) and density anomaly (represented by darkness) of such maxima, if they occur in the lower mantle (depth > 670 km) and exceed 0.125 %, else similar to Fig. 2.

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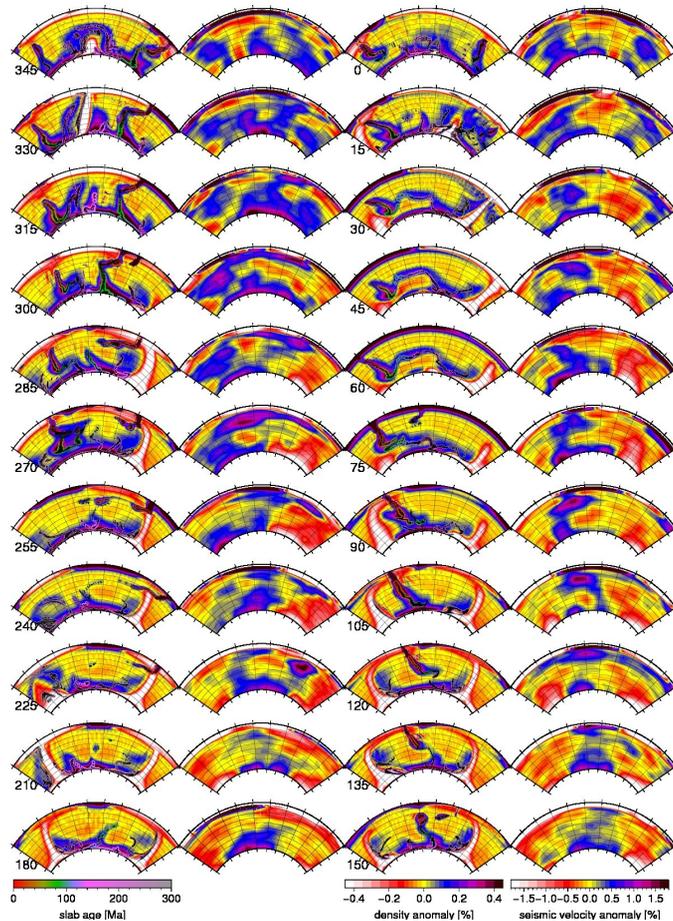
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**Fig. 5.** Caption on next page.

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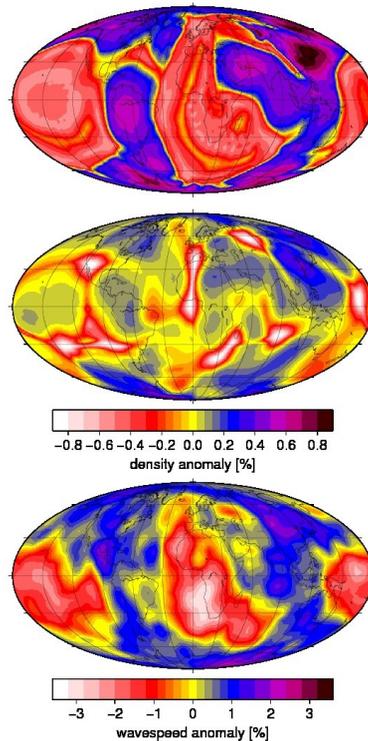
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**Fig. 6.** Map views of predicted thermal density anomalies and seismic wavespeed anomalies in the lowermost 100 km of the mantle. Top panel: density anomalies in the model st12den-2 with a thermochemical layer at the base of the mantle, described in Sect. 2 and shown in Figs. 4 and 5. Center panel: same for a model without thermochemical layer. No phase boundary is considered, and slabs are inserted at depths 600 and 650 km 12 Myr after subduction. However the latter two differences only change results in a minor way; the main difference is due to presence or absence of a thermochemical layer. Bottom panel: composite tomography SMEAN (Becker and Boschi, 2002). Again, mean values are set to zero in all cases.

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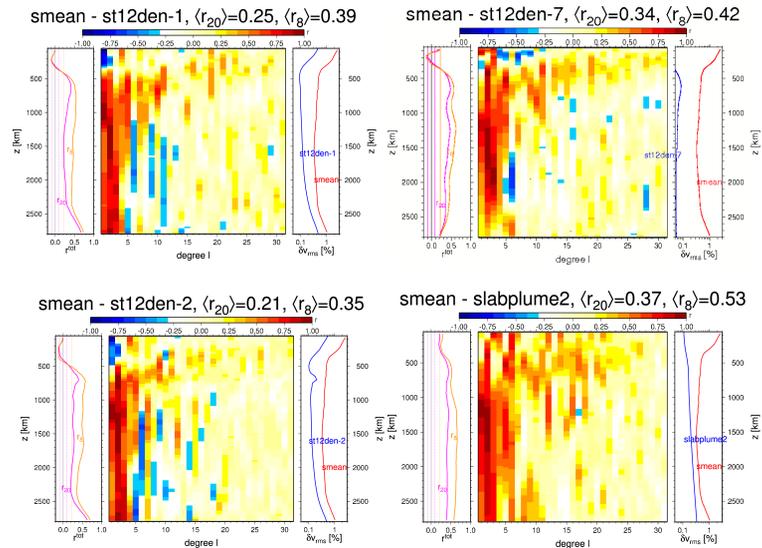
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**Fig. 7.** Correlation between geodynamic models and composite tomographic model SMEAN (Becker and Boschi, 2002): the left parts of each panel give correlations  $r_8$  for all degrees up to  $l = 8$  and  $r_{20}$  for all degrees up to  $l = 20$ , the center parts of each panel correlation as a function of depth and spherical harmonic degree, the right parts of each panel give the RMS anomaly of density (for geodynamic models) or seismic wavespeed (for tomography models) as a function of depth. Top left: model st12den-1, with input of slabs over a wider depth range (and further described in the text). Bottom left: st12den-2, the model shown in most other figures. Top right: st12den-7, the “slabs only” geodynamic model without thermo-chemical piles and without diffusion of heat across the CMB – otherwise same modeling assumptions as st12den-1. Bottom right: a model where tilted mantle plumes with moving source at the CMB according to Boschi et al. (2007) and based on the modeling procedure described in Steinberger and Antretter (2006) have been added to the slab model st12den-7.

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